

## Determination of sulfur speciation in apatites from Martian meteorites (shergottites) using $\mu$ -XANES

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Apatite is a ubiquitous accessory mineral in planetary materials, including in Martian meteorites. Sulfur in terrestrial apatites is mostly  $S^{6+}$  as they are formed in relatively oxidized environments [1]. Recently,  $S^{2-}$ -only bearing apatites have been documented in lunar [1] and terrestrial [2] environments and in experiments [3], the later also reporting simultaneous incorporation of both  $S^{6+}$  and  $S^{2-}$  at intermediate oxygen fugacities ( $fO_2$ ). Thus, it has been suggested that proportions of  $S^{6+}/S^{2-}$  in apatite, together with major element compositions,  $T$  and  $P$ , may record the  $fO_2$  of formation of these apatites [1,3]. Martian rocks, by the virtue of recording intermediate  $fO_2$  between the Moon and Earth [4], may contain  $S^{2-}$ -only or both  $S^{6+}$  and  $S^{2-}$  bearing apatites.

Martian shergottites Shergotty and QUE 94201 record  $fO_2$  of  $\sim IW+1.9$ - $IW+2.8$  and  $\sim IW-1.5$ - $IW-1.0$ , respectively [4], where we expect silicate liquids and the apatites crystallizing from those liquids to contain S primarily as  $S^{2-}$ . To test this hypothesis, we present S-XANES measurements of apatite grains and other associated phases.

Shergotty apatites only show peaks of structural  $S^{2-}$  in apatites (2470 and 2477 eV) with  $S^{6+}/\sum S=0$ . The presence of  $S^{2-}$ -only apatites in Shergotty is consistent with other mineralogical records of  $fO_2$  in this meteorite and suggest that the oxidation state of sulfur records and preserves the  $fO_2$  during igneous crystallization of apatite. QUE 94201 apatites, however, show peaks of both  $S^{2-}$  and  $S^{6+}$  (2481.7 eV), varying in  $S^{6+}/\sum S$  from 0-100%. This meteorite is known to have gone through substantial sulfate-rich alteration and oxidation both on the fusion crust and in the interior upon arrival to Earth's surface [5]. The oxidized nature of apatites, considering together with the low  $fO_2$  recorded by the meteorite and reduced nature of apatites from minimally altered Shergotty and lunar rocks suggest that these measurements and/or the apatite grains themselves are subject to contamination by secondary oxidative alteration events on Earth and/or Mars.

[1] Brounce M. et al. (2019) *Am. Min.* [2] Konecke B.A. et al. (2019) *GCA* [3] Sadove G. et al. (2019) *Ore Geol. Rev.* [4] Wadhwa M. (2001) *Science*. [5] Ross et al. (2010) *LPSC 2010*.